

# Singly-Thermostated Ergodicity in Gibbs' Canonical Ensemble and the 2016 Ian Snook Prize

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## Abstract

For a harmonic oscillator, Nosé's single-thermostat approach to simulating Gibbs' canonical ensemble with dynamics samples only a small fraction of the phase space. Nosé's approach has been improved in a series of three steps: [ 1 ] several two-thermostat sets of motion equations have been found which cover the complete phase space in an ergodic fashion; [ 2 ] sets of single-thermostat motion equations, exerting "weak control" over both forces and momenta, have been shown to be ergodic; and [ 3 ] sets of single-thermostat motion equations exerting weak control over two velocity moments provide ergodic phase-space sampling for the oscillator and for the rigid pendulum, but not for the quartic oscillator or for the Mexican Hat potential. The missing fourth step, motion equations providing ergodic sampling for anharmonic potentials requires a further advance. The 2016 Ian Snook Prize will be awarded to the author(s) of the most interesting original submission addressing the problem of finding ergodic algorithms for Gibbs' canonical ensemble using a single thermostat.

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## I. GIBBS' CANONICAL ENSEMBLE

From Gibbs' 1902 text *Elementary Principles in Statistical Mechanics*, page 183 :

“If a system of a great number of degrees of freedom is microcanonically distributed in phase, any very small part of it may be regarded as canonically distributed.”

Thus J. Willard Gibbs pointed out that the energy states of a “small” system weakly coupled to a larger “heat reservoir” with a temperature  $T$  have a “canonical” distribution :

$$f(q, p) \propto e^{-\mathcal{H}(q,p)/kT} .$$

with the Hamiltonian  $\mathcal{H}(q, p)$  that of the small system. Here  $(q, p)$  represents the set of coordinates and momenta of that system.

“ *Canonical* ” means simplest or prototypical. The heat reservoir coupled to the small system and responsible for the canonical distribution of energies is best pictured as an ideal-gas thermometer characterized by an unchanging kinetic temperature  $T$  . The reservoir gas consists of many small-mass classical particles engaged in a chaotic and ergodic state of thermal and mechanical equilibrium with negligible fluctuations in its temperature and pressure. Equilibrium within this thermometric reservoir is maintained by collisions as is described by Boltzmann's equation. His “H Theorem” establishes the Maxwell-Boltzmann velocity distribution found in the gas. See Steve Brush's 1964 translation of Boltzmann's 1896 text *Vorlesungen über Gastheorie*.

Prior to fast computers texts in statistical mechanics were relatively formal with very few figures and only a handful of numerical results. In its more than 700 pages Tolman's 1938 tome *The Principles of Statistical Mechanics* includes only two Figures. [ The more memorable one, a disk colliding with a triangle, appears on the cover of the Dover reprint volume. ] Today the results-oriented graphics situation is entirely different as a glance inside any recent issue of *Science* confirms.

## II. NOSÉ-HOOVER CANONICAL DYNAMICS – LACK OF ERGODICITY

In 1984, with the advent of fast computers and packaged computer graphics software already past, Shuichi Nosé set himself the task of generalizing molecular dynamics to mimic

Gibbs' canonical distribution<sup>1,2</sup>. In the end his approach was revolutionary. It led to a new form of heat reservoir described by a single degree of freedom with a logarithmic potential, rather than the infinitely-many oscillators or gas particles discussed in textbooks. Although the theory underlying Nosé's approach was cumbersome Hoover soon pointed out a useful simplification<sup>3,4</sup> : Liouville's flow equation in the phase space provides a direct proof that the "Nosé-Hoover" motion equations are consistent with Gibbs' canonical distribution. Here are the motion equations for the simplest interesting system, a single one-dimensional harmonic oscillator :

$$\dot{q} = (p/m) ; \dot{p} = -\kappa q - \zeta p ; \dot{\zeta} = [ (p^2/mkT) - 1 ] / \tau^2 .$$

The "friction coefficient"  $\zeta$  stabilizes the kinetic energy ( $p^2/2m$ ) through integral feedback, extracting or inserting energy as needed to insure a time-averaged value of precisely  $(kT/2)$  . The parameter  $\tau$  is a relaxation time governing the rate of the thermostat's response to thermal fluctuations. In what follows we will set all the parameters and constants ( $m, \kappa, k, T, \tau$ ) equal to unity, purely for convenience. Then the Nosé-Hoover equations have the form :

$$\dot{q} = p ; \dot{p} = -q - \zeta p ; \dot{\zeta} = p^2 - 1 [ \text{NH} ] .$$

Liouville's phase-space flow equation, likewise written here for a single degree of freedom, is just the usual continuity equation for the three-dimensional flow of a probability density in the  $(q, p, \zeta)$  phase space :

$$\dot{f} = (\partial f / \partial t) + \dot{q}(\partial f / \partial q) + \dot{p}(\partial f / \partial p) + \dot{\zeta}(\partial f / \partial \zeta) = -f(\partial \dot{q} / \partial q) - f(\partial \dot{p} / \partial p) - f(\partial \dot{\zeta} / \partial \zeta) .$$

This approach leads directly to the simple [ NH ] dynamics described above. It is easy to verify that Gibbs' canonical distribution needs only to be multiplied by a Gaussian distribution in  $\zeta$  in order to satisfy Liouville's equation.

$$e^{-q^2/2} e^{-p^2/2} e^{-\zeta^2/2} \propto f_{NH} \propto f_G e^{-\zeta^2/2} \longrightarrow (\partial f_{NH} / \partial t) \equiv 0 .$$

Hoover emphasized that the simplest thermostated system, a harmonic oscillator, does *not* fill out the entire Gibbs' distribution in  $(q, p, \zeta)$  space. It is not "ergodic" and fails to reach all of the oscillator phase space. In fact, with *all* of the parameters ( mass, force constant, Boltzmann's constant, temperature, and relaxation time  $\tau$  ) set equal to unity only six percent of the Gaussian distribution is involved in the chaotic sea<sup>5</sup>. See **Figure 1** for a cross section of the Nosé-Hoover sea in the  $p = 0$  plane. The complexity in the figure, where

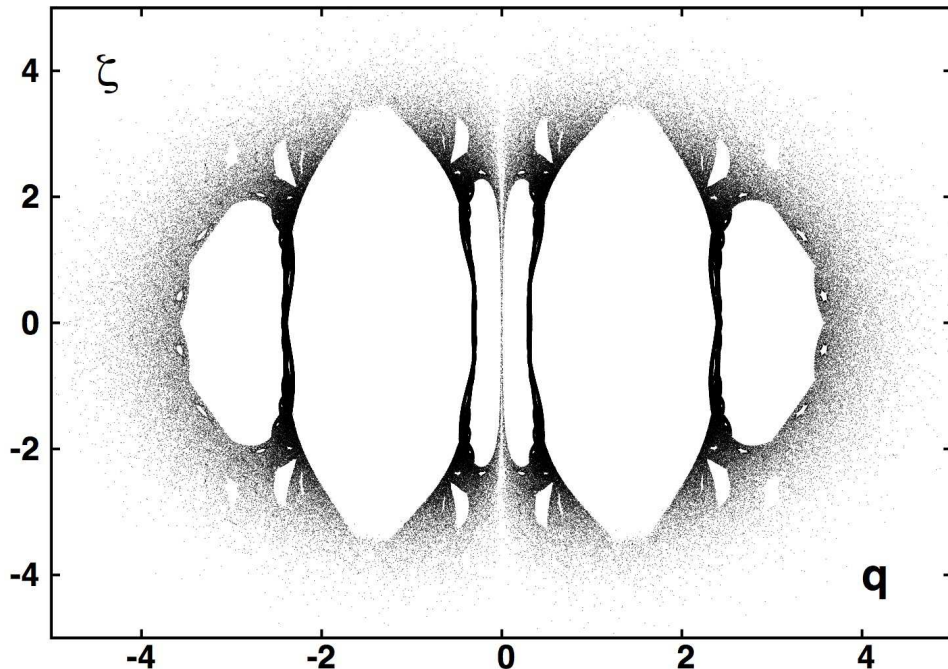


FIG. 1: The  $p = 0$  cross section of the chaotic sea for the Nosé-Hoover harmonic oscillator. 502 924 crossings of the plane are shown. The fourth-order Runge-Kutta integration used a timestep  $dt = 0.0001$ . A point was plotted whenever the product  $p_{old}p_{new}$  was negative.

the “holes” correspond to two-dimensional tori in the three-dimensional  $(q, p, \zeta)$  phase space, is due to the close relationship of the Nosé-Hoover thermostated equations to conventional chaotic Hamiltonian mechanics with its infinitely-many elliptic and hyperbolic points.

### III. MORE GENERAL THERMOSTAT IDEAS

New varieties of thermostats, some of them Hamiltonian and some not, appeared over the ensuing 30-year period following Nosé’s work<sup>6–18</sup>. This list is by no means complete. Though important, simplicity is not the sole motivation for abandoning purely-Hamiltonian thermostats. Relatively recently we pointed out that Hamiltonian thermostats are incapable of generating or absorbing heat flow<sup>6,7</sup>. The close connection between changing phase volume and entropy production guarantees that Hamiltonian mechanics is fundamentally

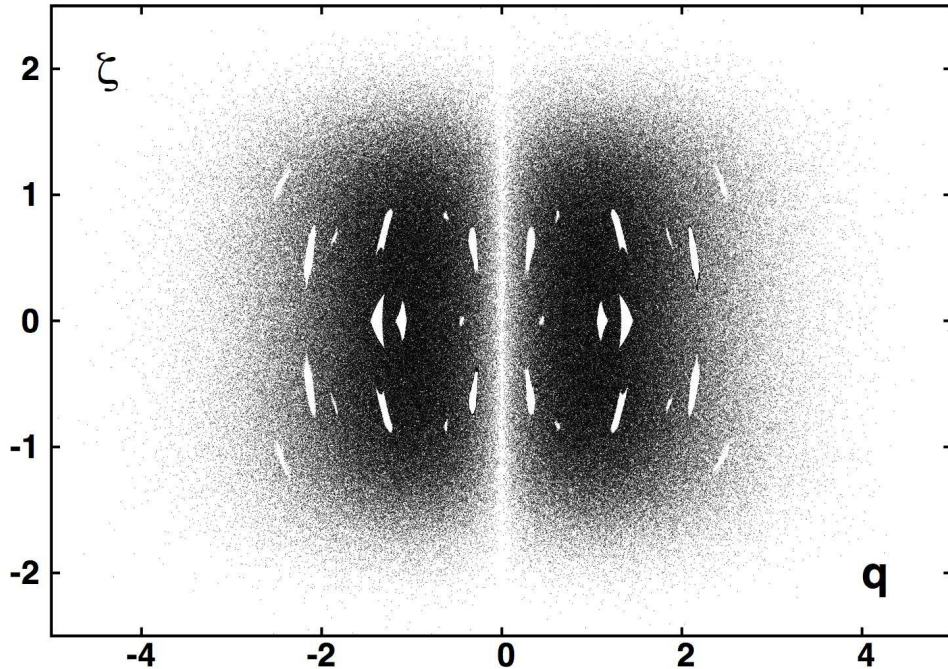


FIG. 2: The  $p = 0$  cross section of the chaotic sea for an oscillator governed by Brańka, Kowalik, and Wojciechowski's choice of the motion equation,  $\ddot{q} = \dot{p} = -q - \zeta^3 p$  ;  $\dot{\zeta} = p^2 - 1$  . 20 billion timesteps, with  $dt = 0.0001$ , resulted in 636 590 crossings of the  $p = 0$  section, using the integration procedure of Figure 1.

inconsistent with irreversible flows.

At equilibrium Brańka, Kowalik, and Wojciechowski<sup>8</sup> followed Bulgac and Kusnezov<sup>9,10</sup> in emphasizing that *cubic* frictional forces,  $-\zeta^3 p$ , which also follow from a novel Hamiltonian, promote a much better coverage of phase space, as shown in **Figure 2** . The many small holes in the  $p = 0$  cross section show that this approach also lacks ergodicity.

### A. Joint Control of Two Velocity Moments

Attempts to improve upon this situation led to a large literature with the most useful contributions applying thermostating ideas with two or more thermostat variables<sup>9,10</sup> . An example, applied to the harmonic oscillator, was tested by Hoover and Holian<sup>11</sup> and found

to provide all of Gibbs' distribution :

$$\dot{q} = p ; \dot{p} = -q - \zeta p - \xi p^3 ; \dot{\zeta} = p^2 - 1 ; \dot{\xi} = p^4 - 3p^2 \text{ [ HH ]}$$

The two thermostat variables  $(\zeta, \xi)$  together guarantee that both the second and the fourth moments of the velocity distribution have their Maxwell-Boltzmann values [ 1 and 3 ] . Notice that two-dimensional cross sections like those in the Figures are no longer useful diagnostics for ergodicity once the phase-space dimensionality exceeds three.

### B. Joint Control of Coordinates and Velocities

In 2014 Patra and Bhattacharya<sup>12</sup> suggested thermostating both the coordinates and the momenta :

$$\dot{q} = p - \xi q ; \dot{p} = -q - \zeta p ; \dot{\zeta} = p^2 - 1 ; \dot{\xi} = q^2 - 1 \text{ [ SEPB ]} .$$

an approach already tried by Sergi and Ezra in 2001<sup>13</sup>.

A slight variation of the Sergi-Ezra-Patra-Bhattacharya thermostat takes into account Bulgac and Kusnezov's observation that cubic terms favor ergodicity :

$$\dot{q} = p - \xi^3 q ; \dot{p} = -q - \zeta p ; \dot{\zeta} = p^2 - 1 ; \dot{\xi} = q^2 - 1 \text{ [ PB}_{\text{var}} \text{ ]} .$$

These last two-thermostat equations appear to be a good candidate for ergodicity, reproducing the second and fourth moments of  $(q, p, \zeta, \xi)$  within a fraction of a percent. We have not carried out the thorough investigation that would be required to establish their ergodicity as the single-thermostat models are not only simpler but also much more easily diagnosed because their sections are two-dimensional rather than three-dimensional.

## IV. SINGLE-THERMOSTAT ERGODICITY

Combining the ideas of "weak control" and the successful simultaneous thermostating of coordinates and momenta<sup>14</sup> led to further trials attempting the weak control of two different kinetic-energy moments<sup>15</sup>. One choice out of the hundreds investigated turned out to be successful for the harmonic oscillator :

$$\dot{q} = p ; \dot{p} = -q - \zeta(0.05p + 0.32p^3) ; \dot{\zeta} = 0.05(p^2 - 1) + 0.32(p^4 - 3p^2) \text{ [ "0532 Model" ]} .$$

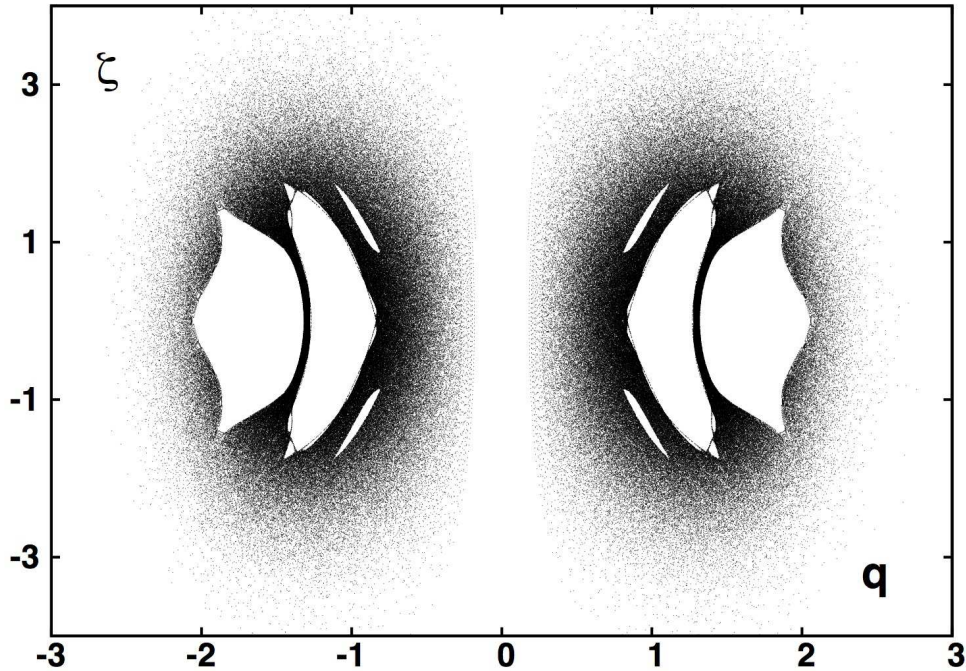


FIG. 3:  $p = 0$  cross section for a singly-thermostated quartic oscillator, with motion equations  $\ddot{q} = \dot{p} = -q^3 - \zeta p^3$  ;  $\dot{\zeta} = p^4 - 3p^2$  . Runge-Kutta integration as in Figures 1 and 2 with 503 709 crossings of the  $p = 0$  plane. Several hundred singly-thermostated attempts failed to obtain canonical ergodicity for the quartic oscillator.

These three oscillator equations passed all of the following tests for ergodicity :

[ 1 ] The moments  $\langle p^2 \rangle = 1$  ;  $\langle p^4 \rangle = 3$  ;  $\langle p^6 \rangle = 15$  were confirmed.

[ 2 ] The independence of the largest Lyapunov exponent to the initial conditions indicated the absence of the toroidal solutions.

[ 3 ] The separation of two nearby trajectories had an average value of 6 :

$$\langle (q_1 - q_2)^2 + (p_1 - p_2)^2 + (\zeta_1 - \zeta_2)^2 \rangle = 2 + 2 + 2 = 6 .$$

[ 4 ] The times spent at positive and negative values of  $\{ q, p, \zeta \}$  were close to equal.

[ 5 ] The times spent in regions with each of the 3! orderings of the three dependent variables were equal for long times.

These five criteria were useful tools for confirming ergodicity. Evidently weak control is the key to efficient ergodic thermostating of oscillator problems.



FIG. 4: Shuichi Nosé ( 1951-2005 ) and Ian Snook ( 1945-2013 )

## V. A FLY IN THE OINTMENT, THE QUARTIC POTENTIAL

The success in thermostating the harmonic oscillator led to like results for the simple pendulum but *not* for the quartic potential<sup>15</sup>. See **Figure 3**. This somewhat surprising setback motivates the need for more work and is the subject of the Ian Snook Prize for 2016. This Prize will be awarded to the author(s) of the most interesting original work exploring the ergodicity of single-thermostated statistical-mechanical systems. The systems are not at all limited to the examples of the quartic oscillator and the Mexican Hat potential but are left to the imagination and creativity of those entering the competition.

## VI. CONCLUSIONS – IAN SNOOK PRIZE FOR 2016

It is our intention to reward the most interesting and convincing entry submitted for publication to Computational Methods in Science and Technology ( [www.cmst.eu](http://www.cmst.eu) ) prior to 31 January 2017. The 2016 Ian Snook prize of \$500 dollars will be presented to the winner in early 2017. An Additional Prize of the same amount will likewise be presented by the Institute of Bioorganic Chemistry of the Polish Academy of Sciences ( Poznan Supercomputing and Networking Center ). We are grateful for your contributions. This work is dedicated to the memories of our colleagues, Ian Snook ( 1945-2013 ) and Shuichi Nosé ( 1951-2005 ), shown in **Figure 4** .



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